

# Symmetry Breaking and Phase Transition in Nuclei

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Basic properties of nucleus

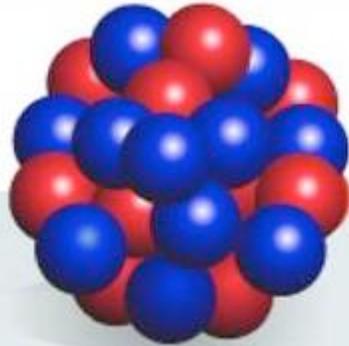
Nucleus in different time scales

Spontaneous Symmetry Breaking in finite time

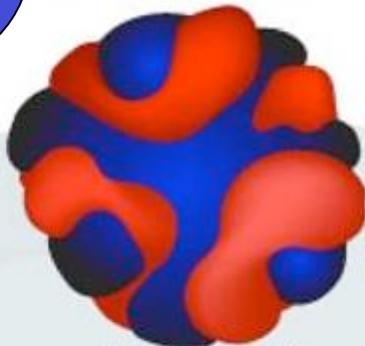
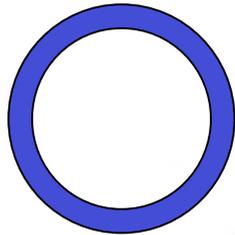
New type of SSB in superdeformed nuclei

Opportunities by RIKEN RI Beam Factory

# Image of nucleus



protons, neutrons



nucleonic densities  
and currents



# Nucleus as a liquid drop

Bethe-Weizsäcker mass formula

$$B(N, Z) = a_V A - a_S A^{2/3} - a_{sym} \frac{(N - Z)^2}{A} - a_C \frac{Z^2}{A^{1/3}} + \delta(A)$$

$$A = N + Z$$

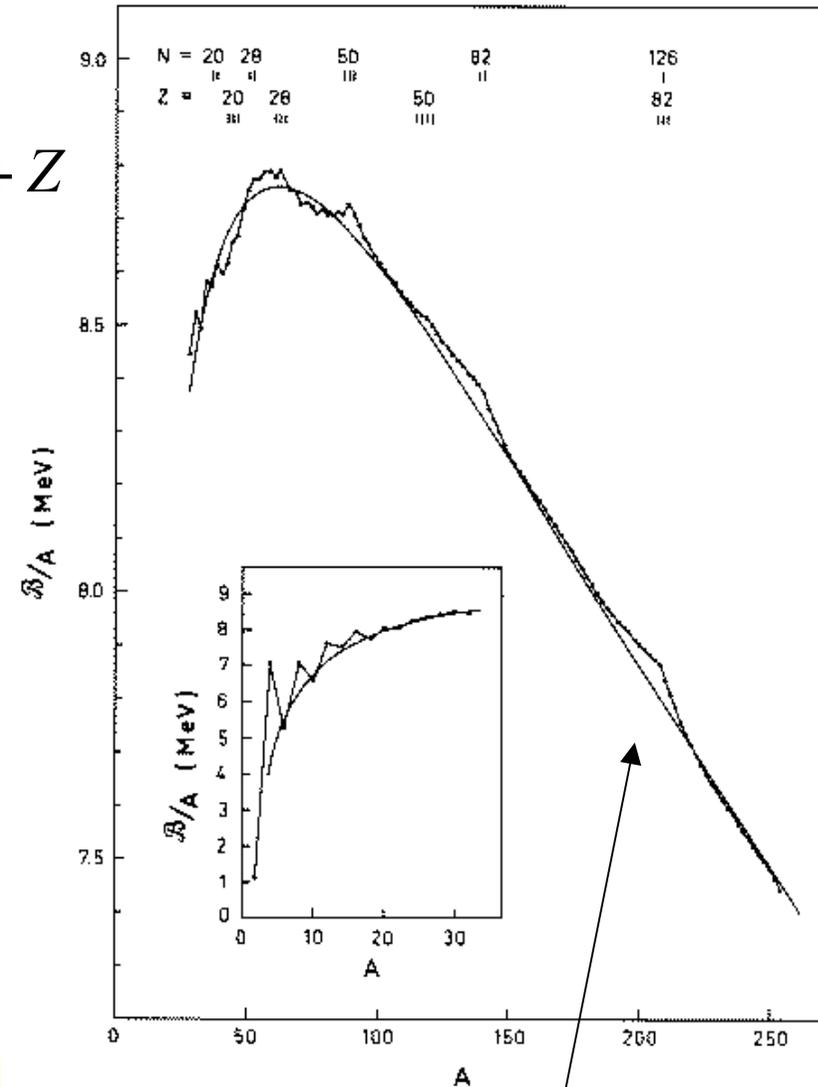
Binding energy per nucleon

$$B/A \approx 8 \text{ MeV}$$

Density

$$\rho \approx 0.14 \text{ fm}^{-3}$$

Nuclear Saturation



*These tiny differences lead to SSB*

# Nucleus as a quantum liquid

- Classical vs Quantum
  - Strength of interaction vs Zero-point kinetic energy

$$V_0 \quad \text{vs} \quad \frac{\hbar^2}{2Mc^2}$$

$c$  : Length scale of the interaction

$V_0$  : Energy scale of the interaction

# Nuclear force vs molecular force

Bohr, Mottelson, Nucl. Str. Vol.1

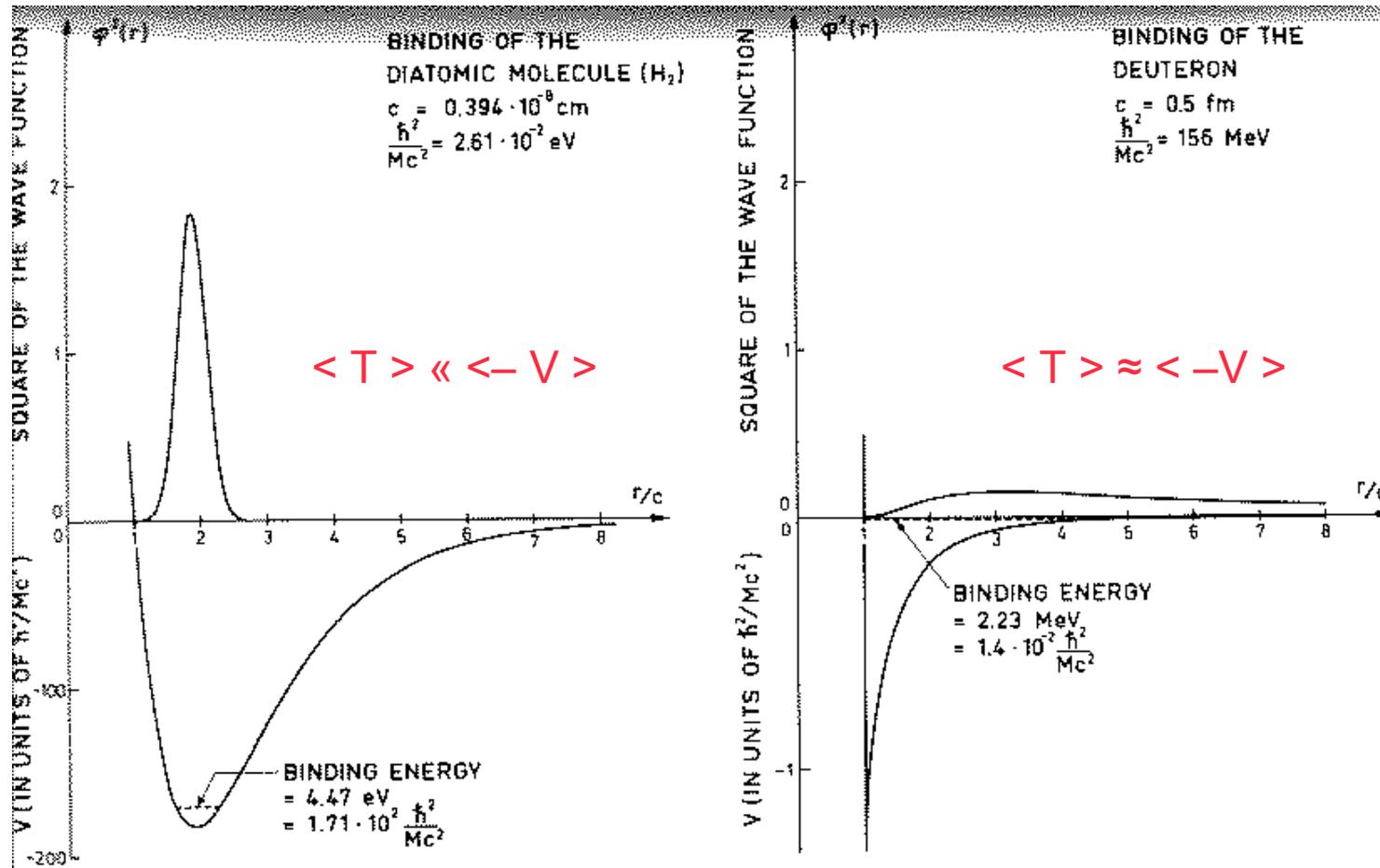


Figure 2-36 The molecular interaction corresponds to a "Morse potential"  $V(r) = D[1 - \exp(-\alpha(r - r_0))]^2 - D$  with the constants adjusted.

Crystallized at low temperature

Classical  $\rightarrow$  MD

Liquid at low temperature

Quantum

# Spontaneous symmetry breaking (SSB) of the rotational and gauge symmetry

- Hamiltonian is rotationally invariant

$$[H, \vec{J}] = 0$$

– **But many nuclei are deformed !**

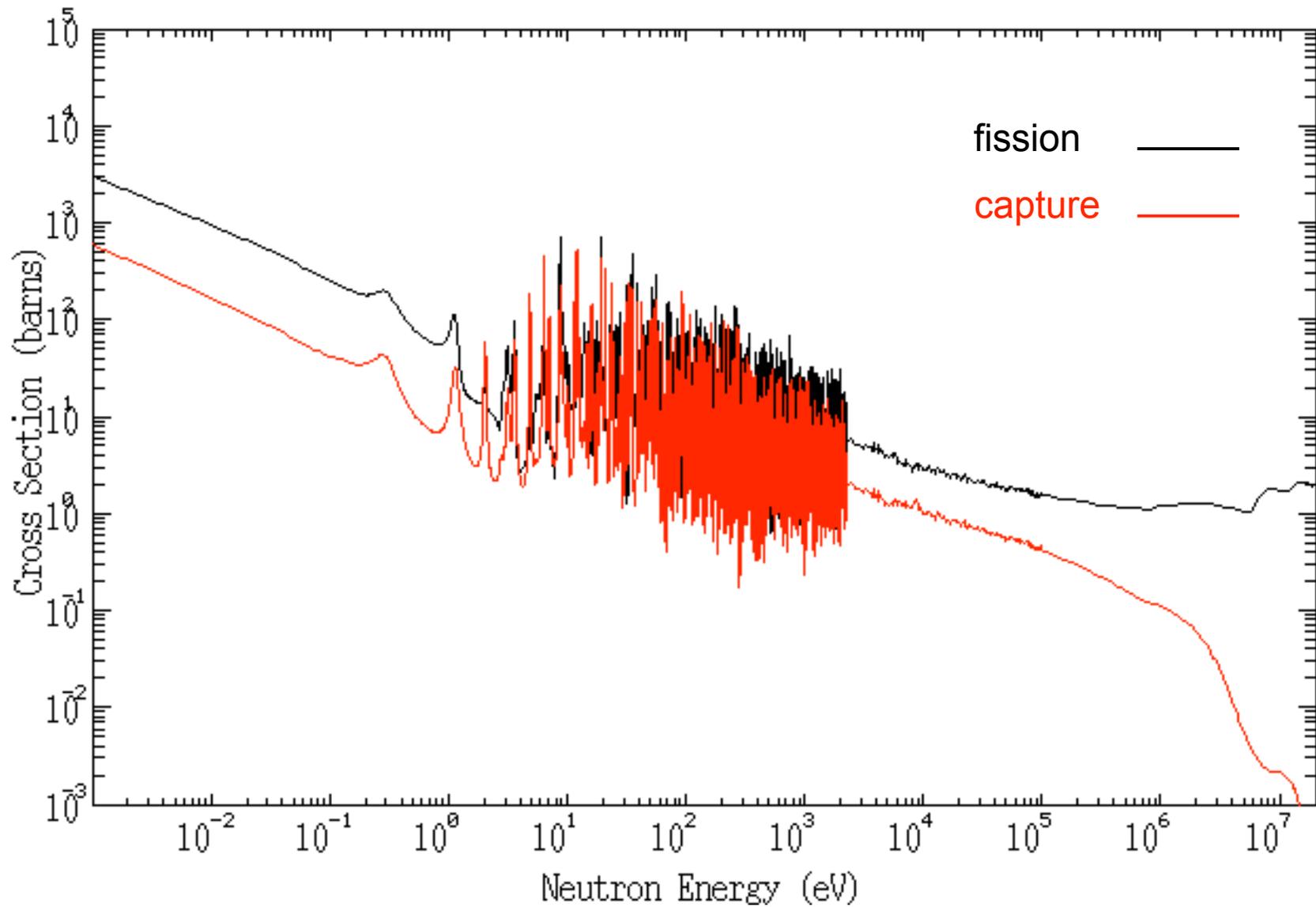
- Hamiltonian is rotationally invariant in the gauge space

$$[H, N] = 0$$

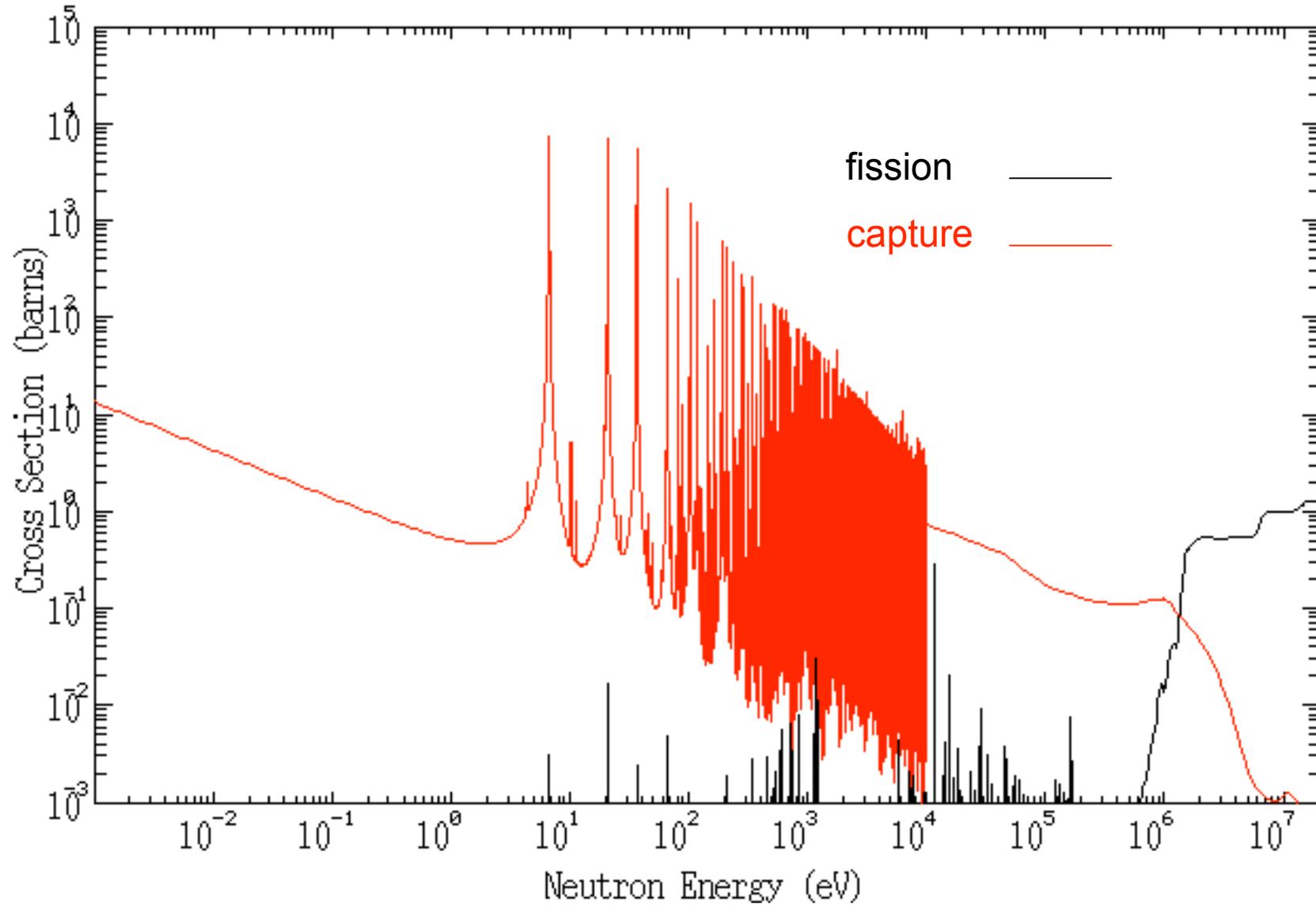
– **But many (heavy) nuclei are superfluid !**

- e. g., Uranium nucleus (235,238) are believed to be deformed and in the superfluid phase.

# Neutron cross section of $^{235}\text{U}$



# Neutron cross section of $^{238}\text{U}$



# Different reaction rate

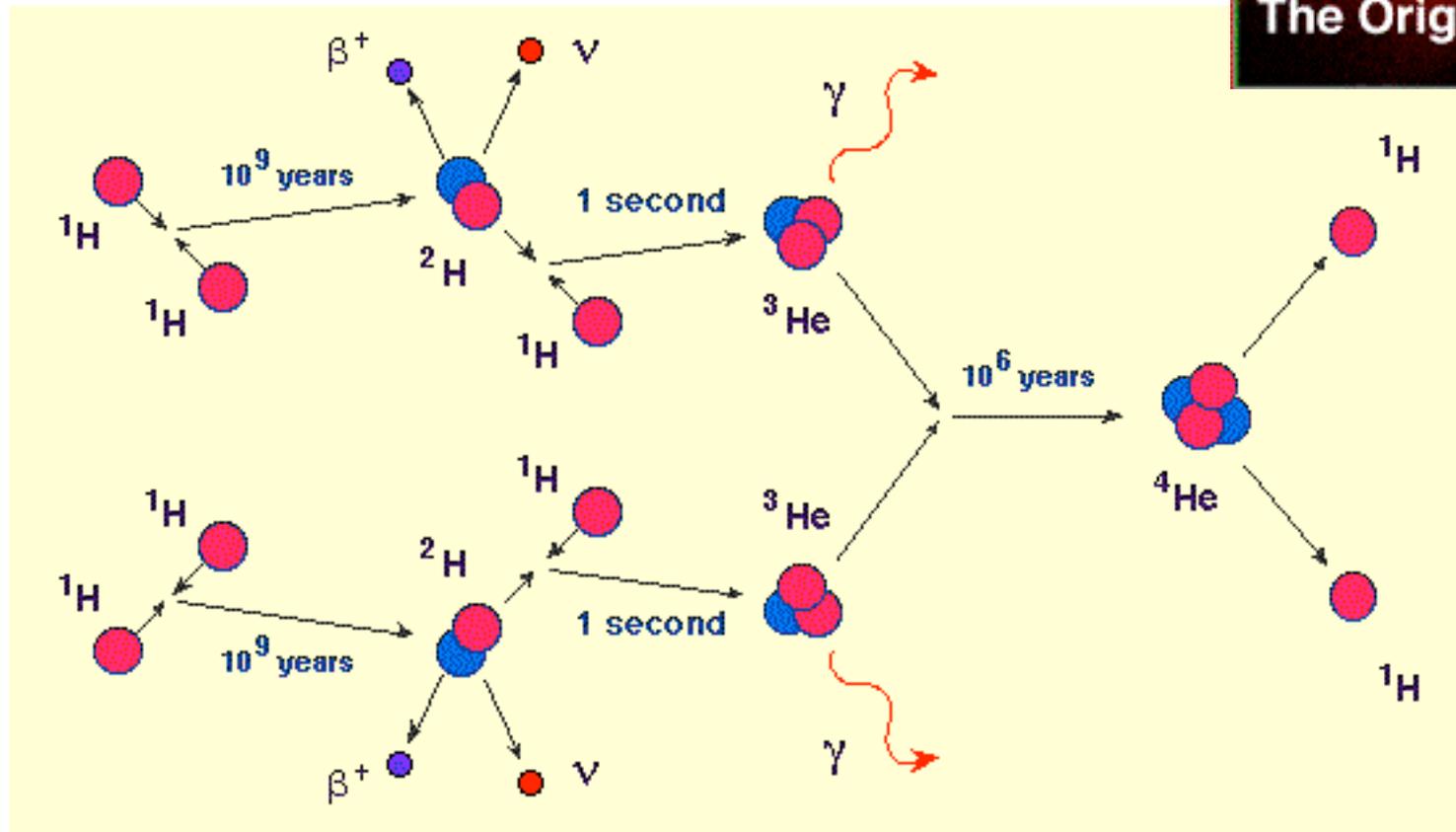
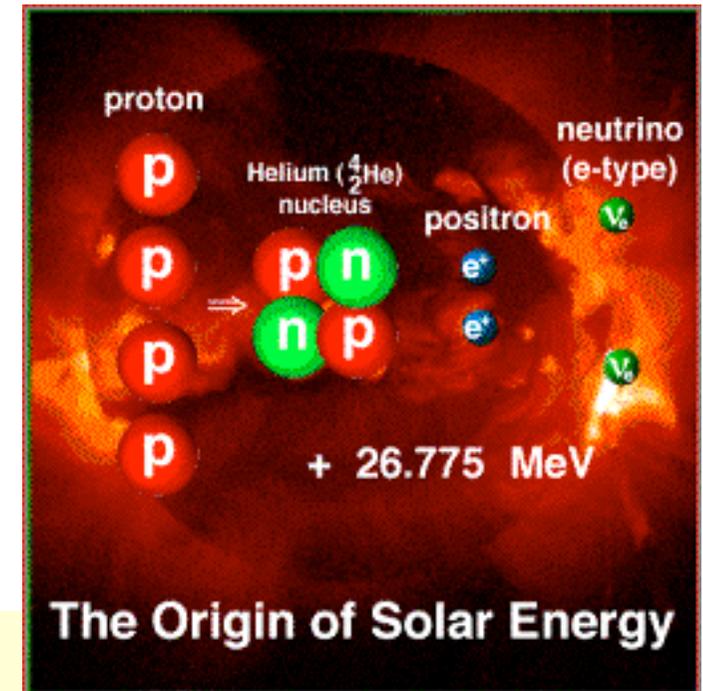
- Transfer reaction (Strong interaction)
  - $^{15}\text{N}(p,\alpha)^{12}\text{C}$
  - $\sigma \sim 0.5 \text{ b}$  ( $E=2 \text{ MeV}$ )
- Capture reaction (Electromagnetic interaction)
  - $^3\text{He}(\alpha,\gamma)^7\text{Be}$
  - $\sigma \sim 10^{-6} \text{ b}$  ( $E=2 \text{ MeV}$ )
- Weak process (Weak interaction)
  - $p(p,e^+\nu)d$
  - $\sigma \sim 10^{-20} \text{ b}$  ( $E=2 \text{ MeV}$ )



*Different time scale*

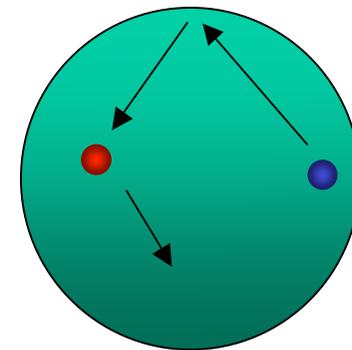
# pp chain (I)

- $p(p, e^+ \nu)d$  reaction determines the lifetime of the sun



# Nucleus with Different Time Scales

- Time period of nucleonic Fermi motion
  - $\tau_F \sim R/v_F \sim 10^{-22}$  sec
- Collision time
  - $\tau_c \gg \tau_F$  (Nucleon near Fermi energy)
  - $\tau_c \geq \tau_F$  (Thermal neutron)

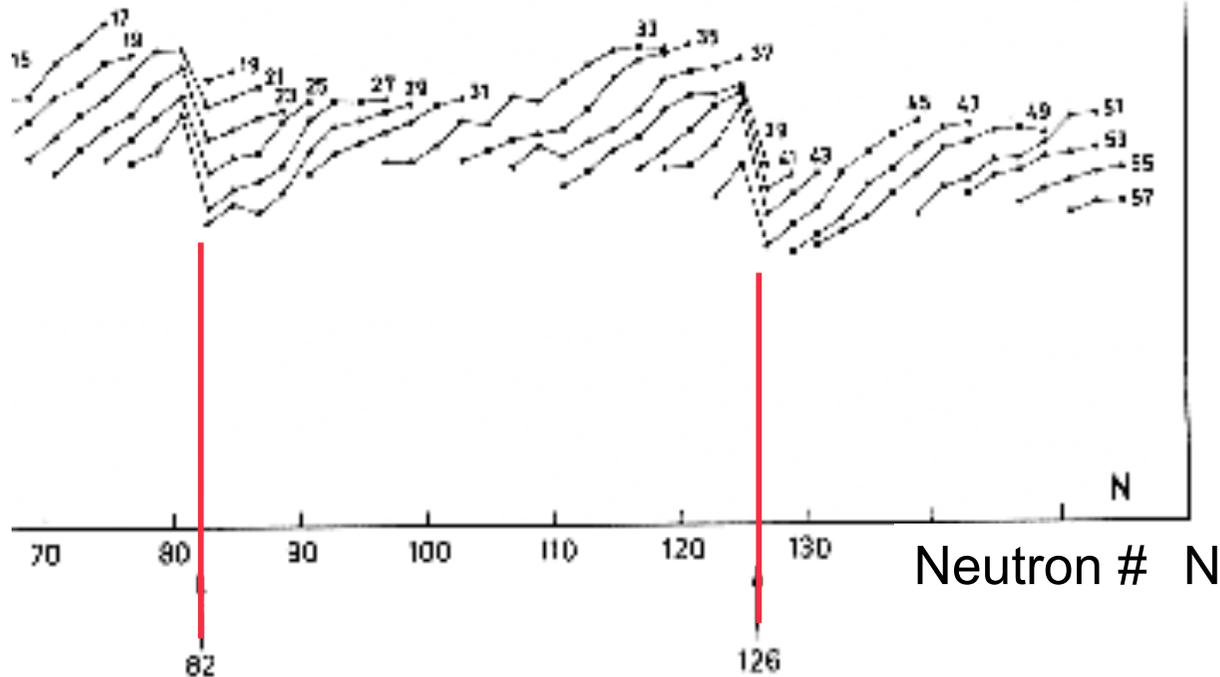


# Nucleonic single-particle motion in nucleus

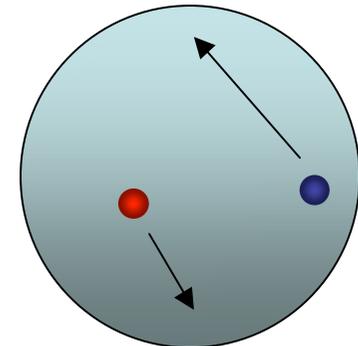
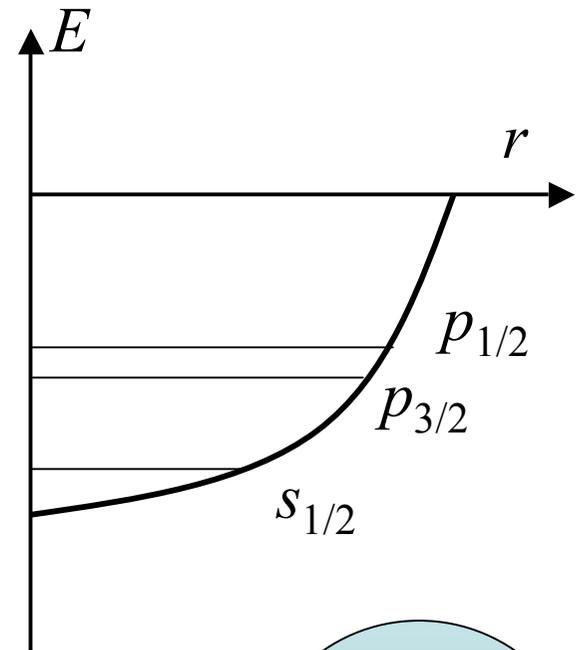
Bohr & Mottelson, Nuclear Structure Vol.1

## Neutron Separation energy

$$S_n(N, Z) = \mathcal{B}(N, Z) - \mathcal{B}(N-1, Z) \quad \begin{array}{l} N \text{ odd} \\ Z \text{ even} \end{array}$$



## Shell Model (Mayer-Jensen)



# SSB in the Nuclear Time Scale

- Time period of nucleonic Fermi motion

- $\tau_F \sim R/v_F \sim 10^{-22}$  sec

- Collision time

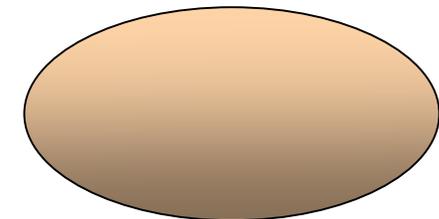
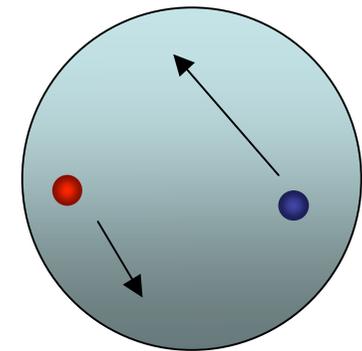
- $\tau_C \gg \tau_F$  (Nucleon near Fermi energy)

- $\tau_C \geq \tau_F$  (Thermal neutron)

- Time scale of symmetry breaking

- $\tau_{SB} \gg \tau_F$  ( $10^{-20}$  sec)

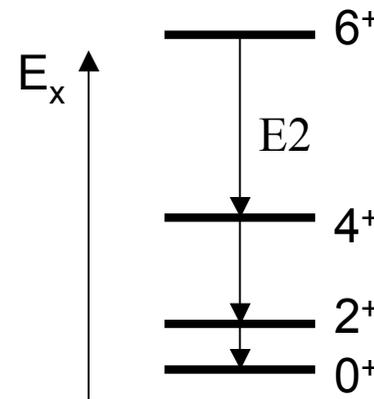
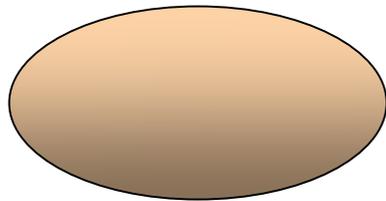
- Broken symmetry is restored by Nambu-Goldstone modes in time of  $\tau_{SB}$



# Nuclear Deformation

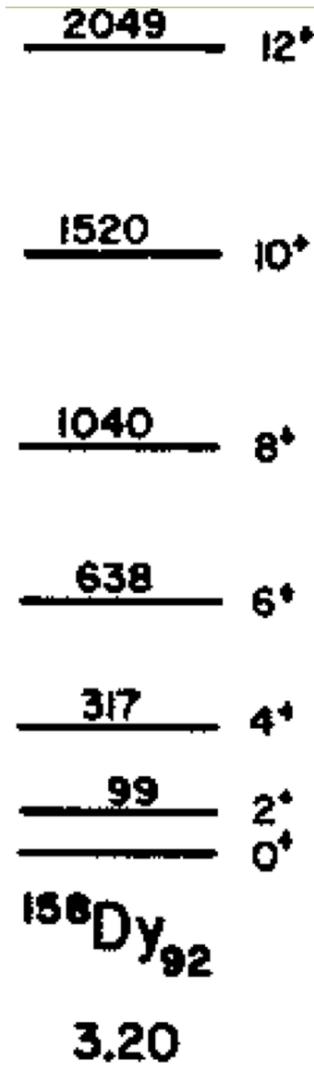
- Rotational excitation spectra
- Strong transition strength

$$E_x = \frac{I(I+1)}{2\mathfrak{I}}$$



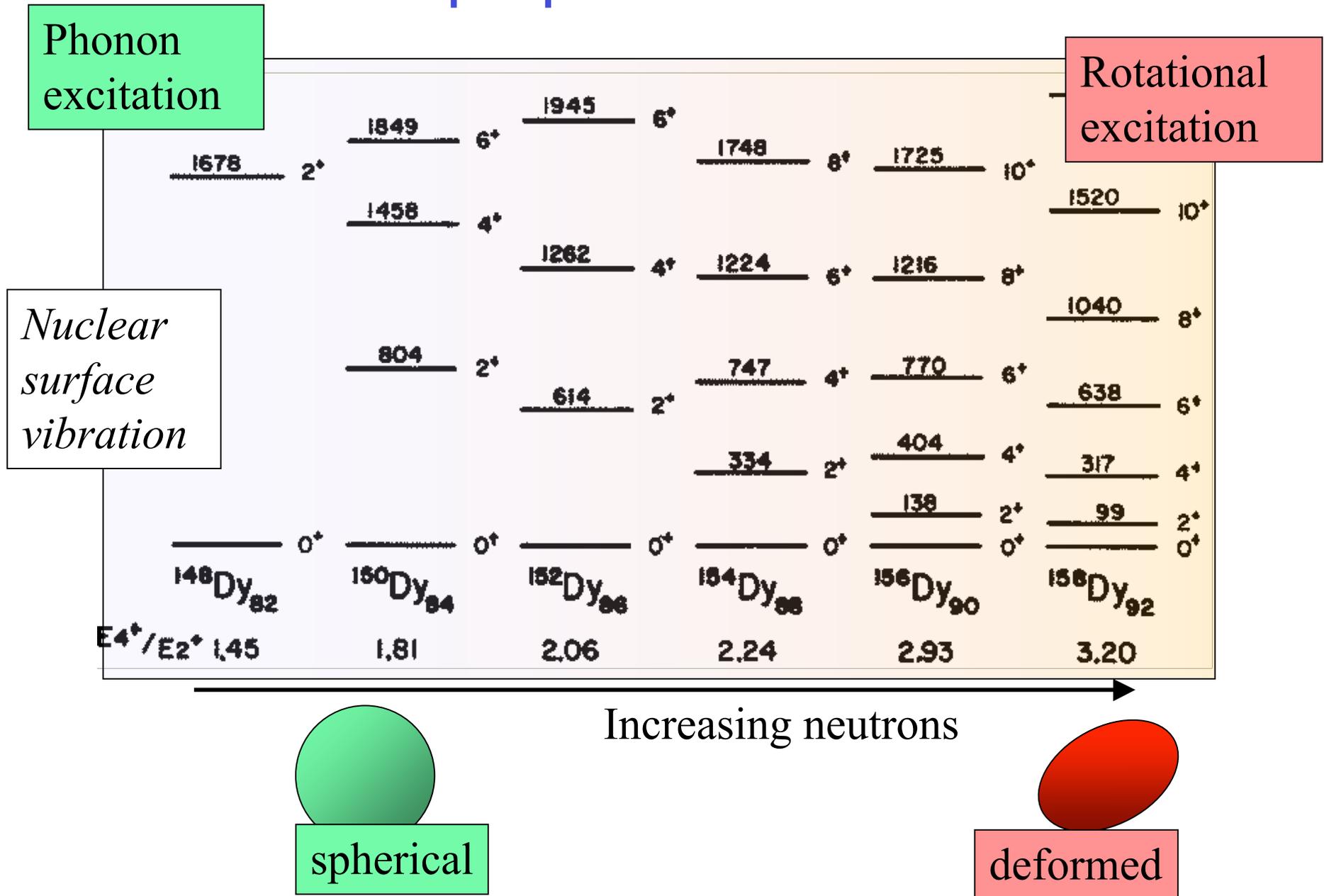
$$B(E2; I \rightarrow I - 2) \equiv \frac{1}{2I+1} \left| \langle I \| M(E2) \| I - 2 \rangle \right|^2$$

More than hundred times larger than typical single-particle values





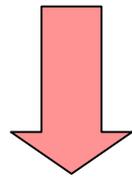
# Shape phase transition



# Nuclear deformation as phonon condensation

$|\Phi_0\rangle$  Ground state

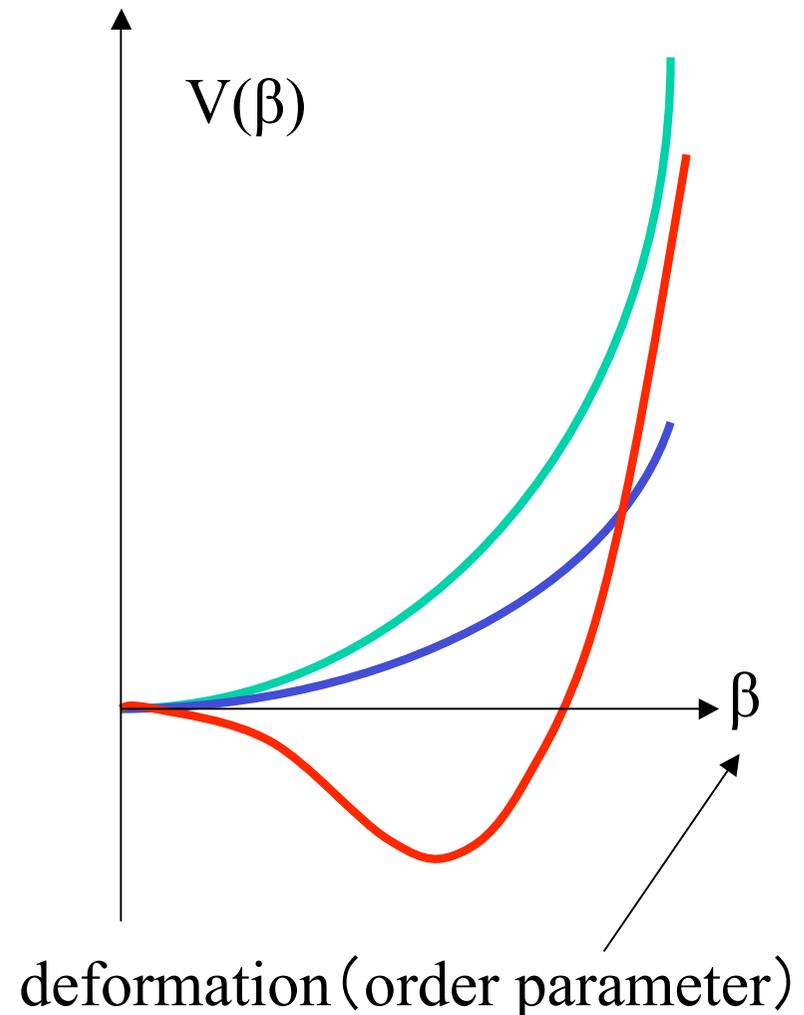
$B_{2^+}^+ |\Phi_0\rangle$  Excited state ( $2^+$ )



Away from the  
closed-shell

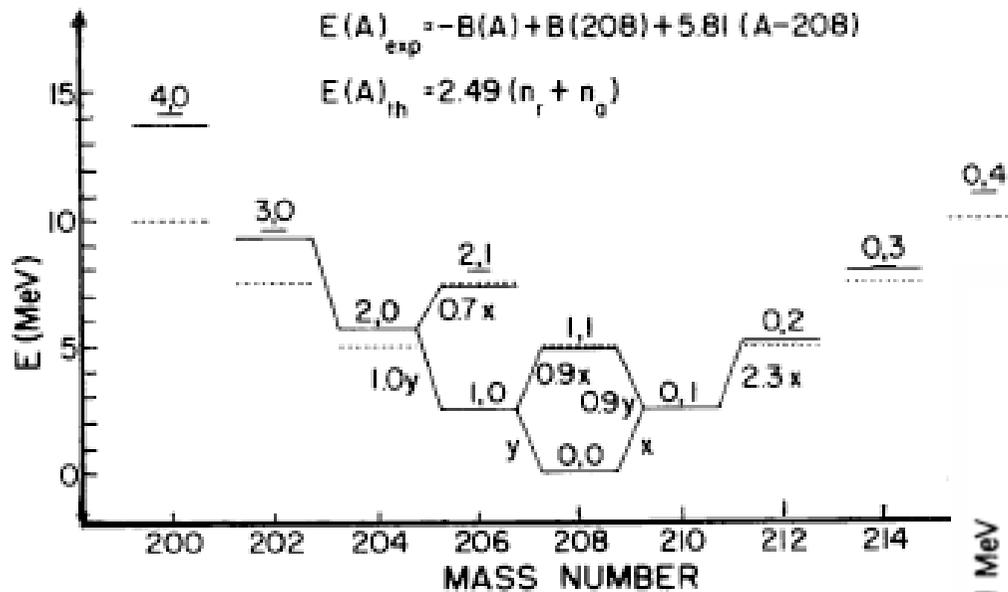
$$|\Phi_{\text{def}}\rangle = \exp(cB_{2^+}^+) |\Phi_0\rangle$$

$$\beta \equiv \langle \Phi_{\text{def}} | r^2 Y_{20}(\hat{r}) | \Phi_{\text{def}} \rangle$$



# Pairing vibration to Pairing rotation

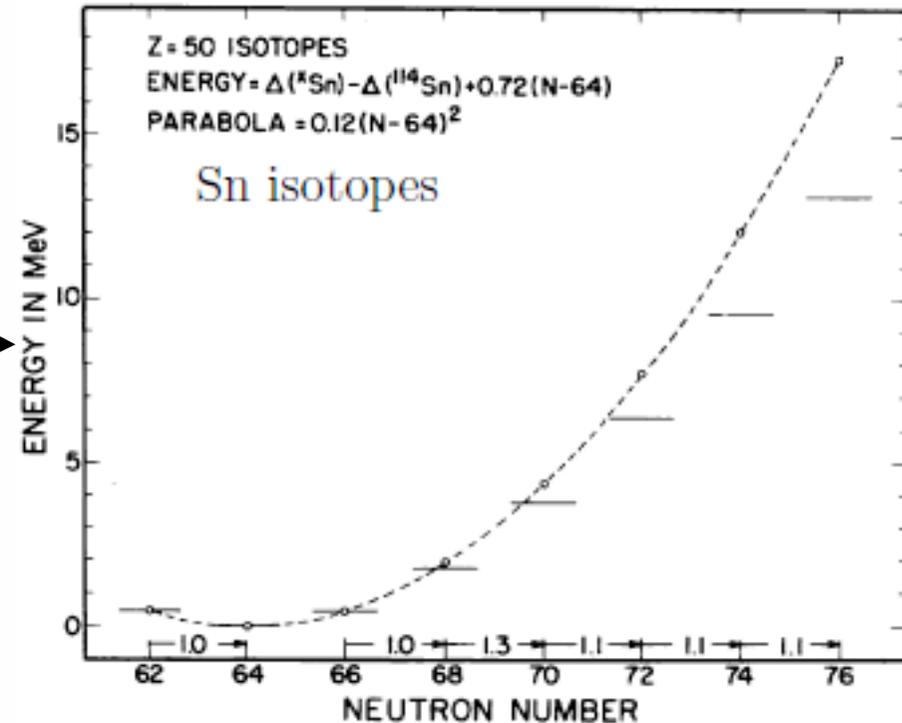
Pb isotopes



$$E_N = \frac{(N - N_0)^2}{2\mathfrak{I}}$$

← Pair removal      Pair addition →

$$E_N = \hbar\omega |N - N_0|$$



# Nuclear superfluidity as pair condensation

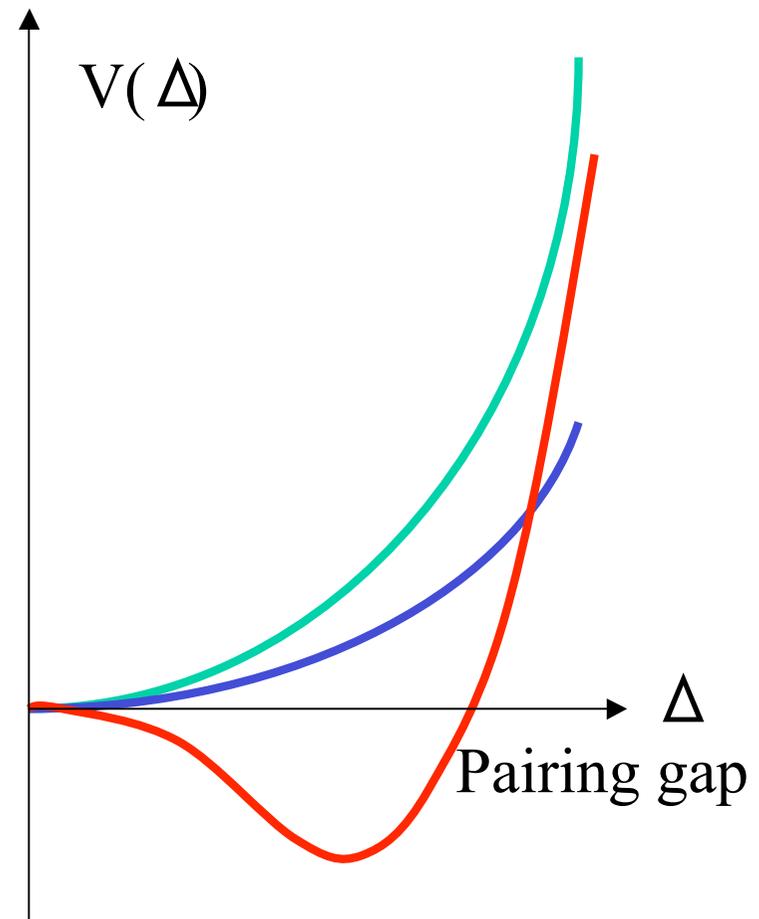
$|\Phi_0\rangle$  Ground state

$P_{0^+}^+ |\Phi_0\rangle$  Pairing vibration

↓ Away from the closed-shell

$$|\Phi_{\text{super}}\rangle = \exp(cP_{0^+}^+) |\Phi_0\rangle$$

$$\Delta \propto \langle \Phi_{\text{super}} | (\hat{\psi}\hat{\psi})_{0^+} | \Phi_{\text{super}} \rangle$$



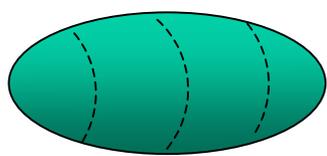
# Symmetry breaking in nuclei

$$e^{i\phi J} |\Psi\rangle \neq |\Psi\rangle$$

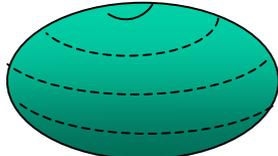
$$e^{i\phi N} |\Psi\rangle \neq |\Psi\rangle$$

Quadrupole deformation

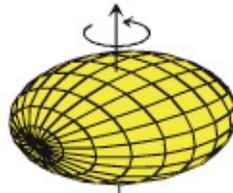
$$\beta_{2\mu} = \langle \Psi | r^2 Y_{2\mu} | \Psi \rangle$$



prolate



oblate



triaxial

Pairing deformation  
(superfluidity)

$$\Delta = \langle \Psi | \hat{\psi}^+ \hat{\psi} | \Psi \rangle$$

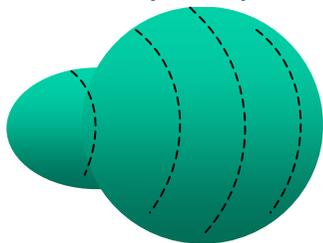
Deformation in the gauge space

*Nuclear Superconductivity*

*Nuclear Superfluidity*

Octupole deformation

$$\beta_{30} = \langle \Psi | r^3 Y_{30} | \Psi \rangle$$



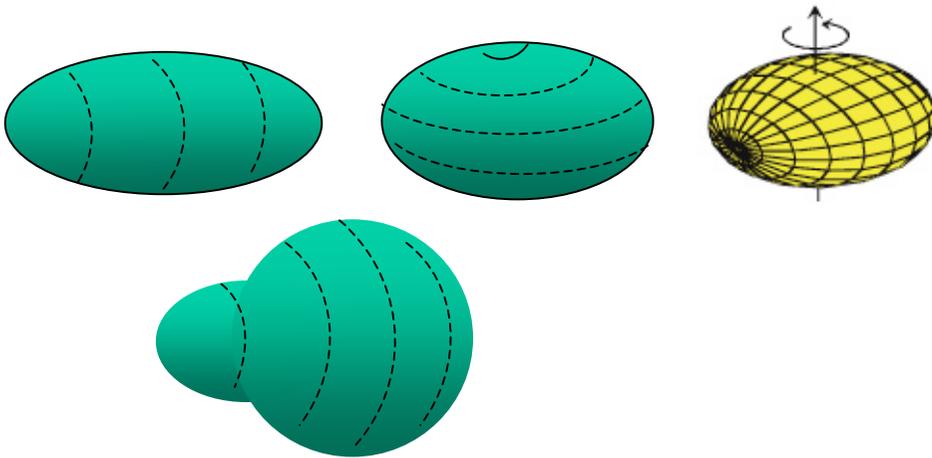
Pear shape ( $\mu=0$ )

$$\hat{P} |\Psi\rangle \neq \pm |\Psi\rangle$$

# Symmetry breaking in nuclei

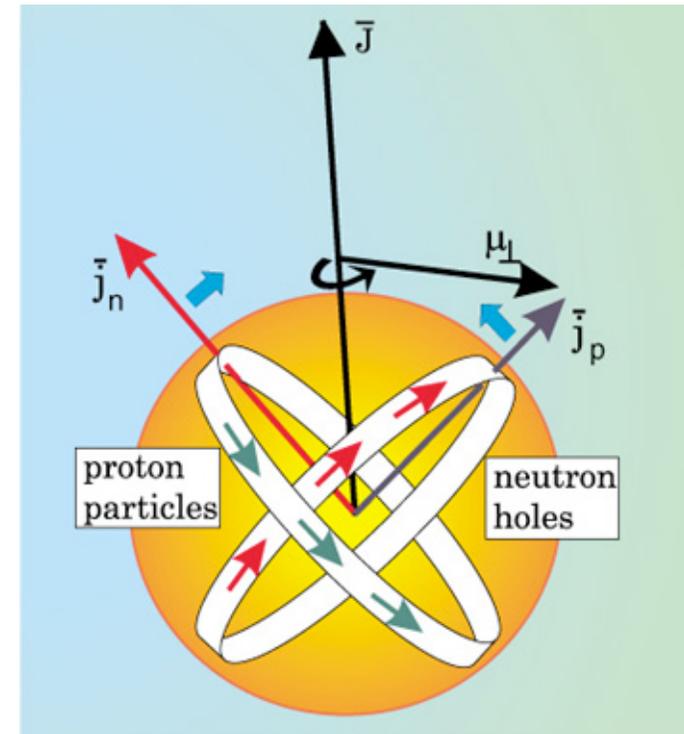
$$e^{i\phi J} |\Psi\rangle \neq |\Psi\rangle$$

Spatial deformation  
(Electric deformation)



S. Frauendorf

Magnetic deformation ?

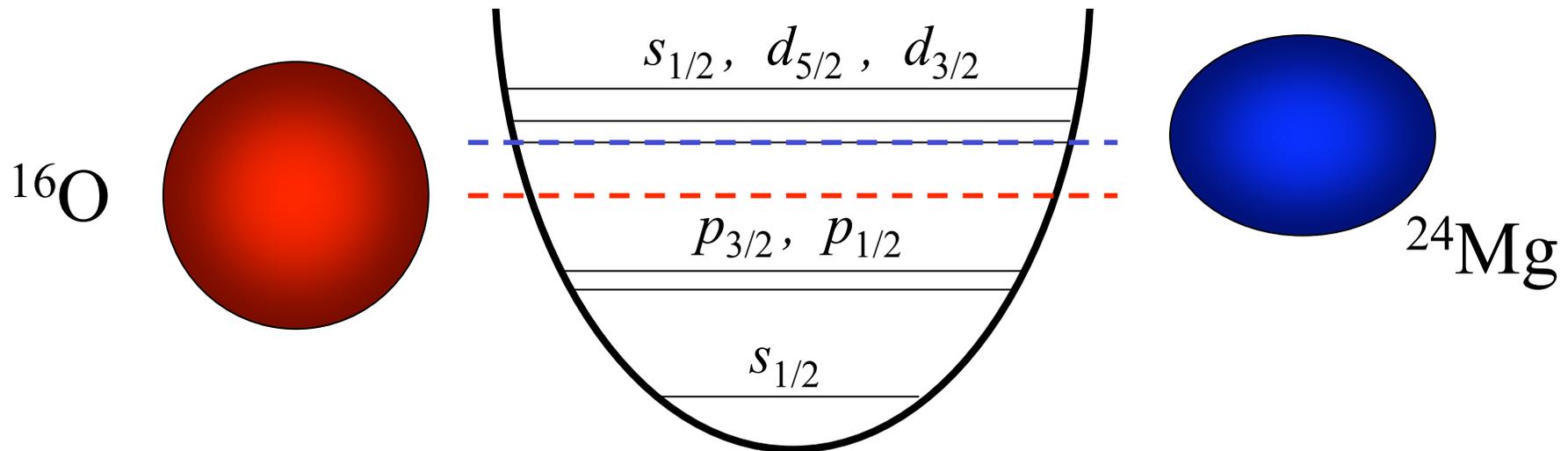


# Shell Structure

Approximate degeneracy of single-particle levels



*Stability* and *Instability* of spherical shape



Driving force of SSB

Jahn-Teller Mechanism

# Shell Structure

Bohr, Mottelson, Nucl. Str. Vol.2

Single-particle energies in a spherical potential can be specified by the quantum numbers  $(n, l)$

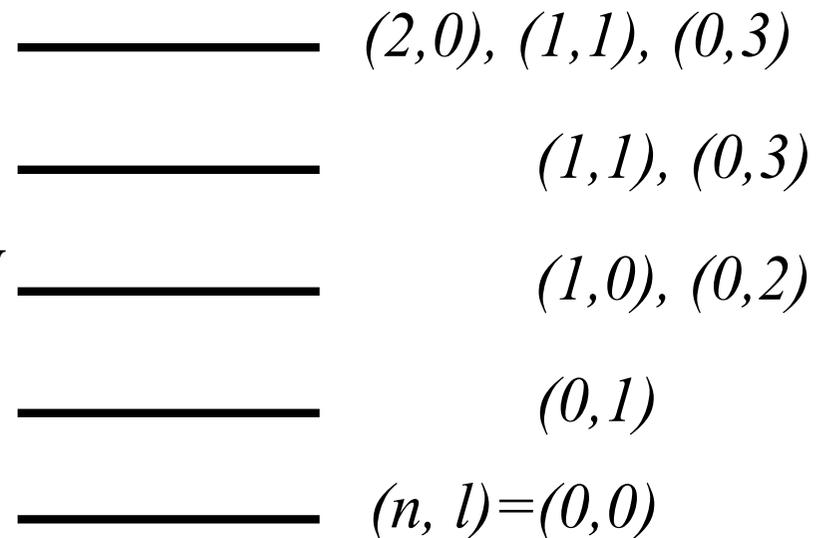
$$\varepsilon(n, l) \approx \varepsilon(n_0, l_0) + \left. \frac{\partial \varepsilon}{\partial n} \right|_0 (n - n_0) + \left. \frac{\partial \varepsilon}{\partial l} \right|_0 (l - l_0)$$

$$\left. \frac{\partial \varepsilon}{\partial n} \right|_0 : \left. \frac{\partial \varepsilon}{\partial l} \right|_0 = a : b$$

Integer  $a:b$   Degeneracy

*Nuclear shell structure*

$$a : b = 2 : 1$$



# Classical correspondence

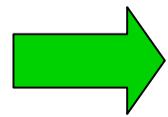
Action-angle variables:  $(l, \varphi)$

$$\frac{\partial \varepsilon}{\partial l} = \varphi$$

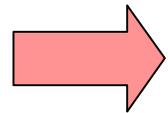
$$\left. \frac{\partial \varepsilon}{\partial n} \right|_0 : \left. \frac{\partial \varepsilon}{\partial l} \right|_0 = \varphi_r : \varphi_l = a : b$$

*Closed periodic orbital*

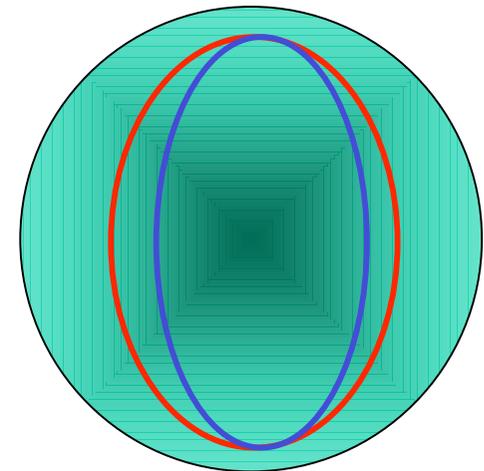
$a : b = 2 : 1$



Elliptic orbits



SSB to ellipsoidal shape  
in open-shell nuclei



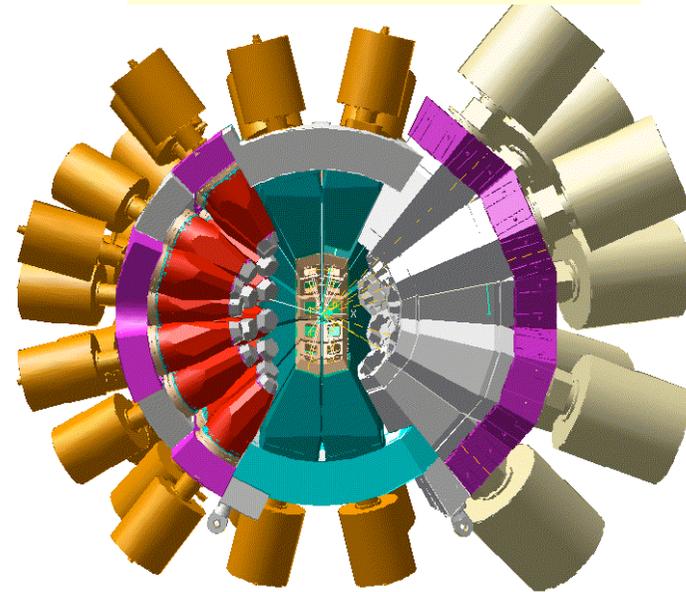
Spatial shape of w.f. created by superposing degenerate s.p. states.

*Germanium gamma-ray detector  
arrays developed in 1980'-90's*

GAMMASPHERE



EUROBALL

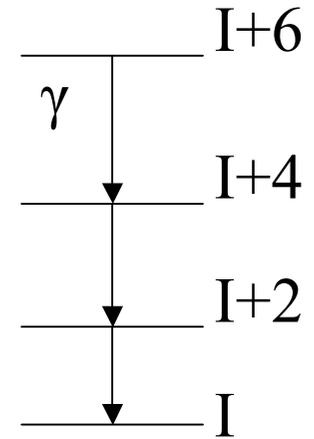


# Discovery of Superdeformed band

1986 @Daresbury, UK

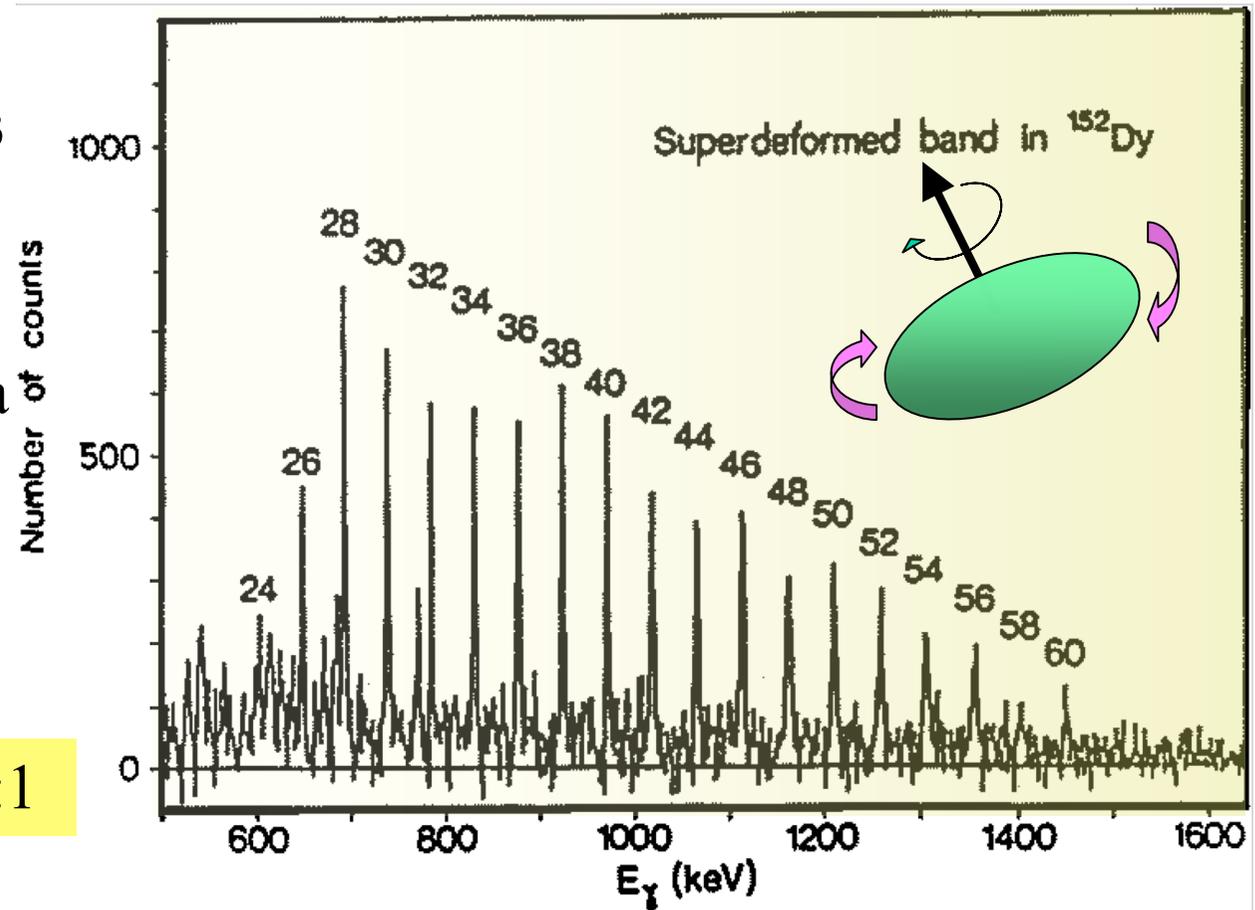
P.J. Twin et al, PRL **57**  
(1986) 811

J.D. Garret et al, Nature **323**  
(1986) 395.

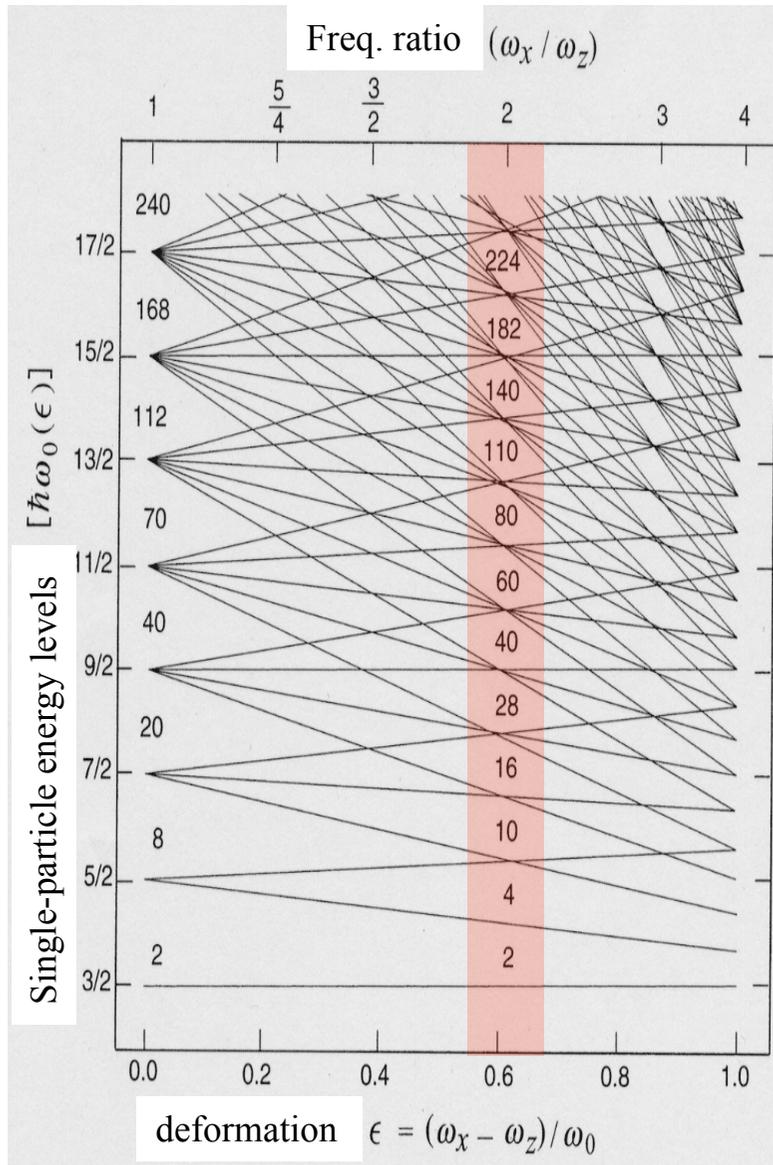


- Large moment of inertia
- Large intraband  $B(E2)$   
 $B(E2) \approx 2000$  W.u.

Major : Minor axes  $\sim 2:1$



# Mechanism of stabilizing superdeformation



3D Harmonic Oscillator Potential

Degeneracy at spherical shape

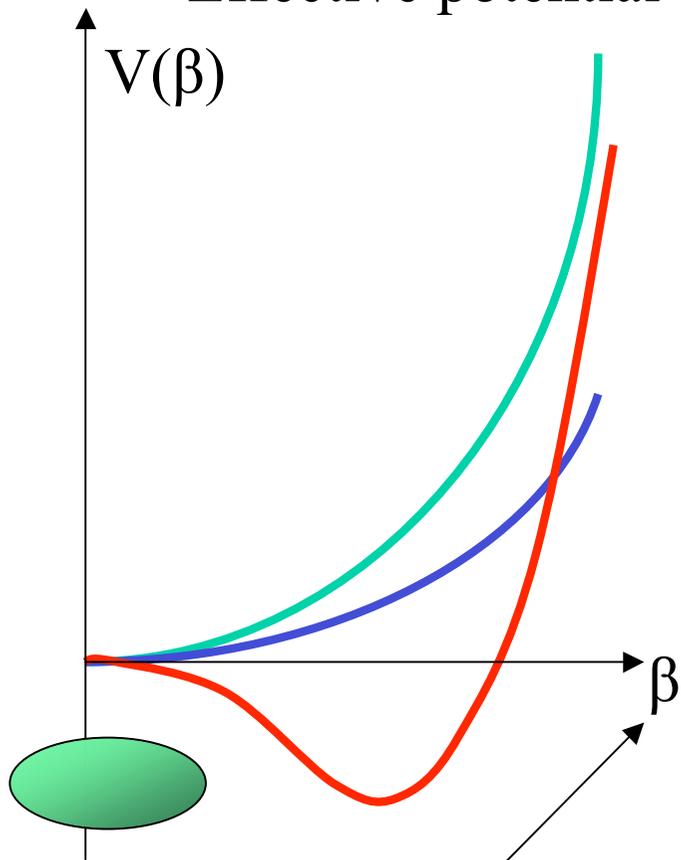
Magic numbers: 2, 8, 20, ...

Degeneracy at 2:1 deformation

Superdeformed magic number:  
2, 4, 10, 16, ...

# What happens if we add neutrons to “closed-shell” superdeformed nuclei?

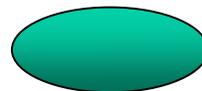
Effective potential



What kind of parameter?

		1945	6*		2049	12*		
1678	2*	1849	6*	1748	8*	1725	10*	
		1458	4*			1520	10*	
				1262	4*	1224	6*	
						1216	8*	
						1040	8*	
		804	2*					
				614	2*	747	4*	
						770	6*	
							638	6*
						334	2*	
						404	4*	
							317	4*
						138	2*	
							99	2*
								0*
0*	0*	0*	0*	0*	0*	0*	0*	
$^{148}\text{Dy}_{82}$	$^{150}\text{Dy}_{84}$	$^{152}\text{Dy}_{86}$	$^{154}\text{Dy}_{88}$	$^{156}\text{Dy}_{90}$	$^{158}\text{Dy}_{92}$			
$E_{4^+}/E_{2^+}$ 1.45	1.81	2.06	2.24	2.93	3.20			

SD  $^{152}\text{Dy}$



$^{162}\text{Dy}$

??

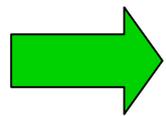
# Typical classical periodic orbits in superdeformed potential

Deformed HO potential

$$(n, l) \Rightarrow (n_x, n_y, n_z)$$

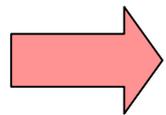
$$\left. \frac{\partial \varepsilon}{\partial n_x} \right|_0 : \left. \frac{\partial \varepsilon}{\partial n_y} \right|_0 : \left. \frac{\partial \varepsilon}{\partial n_z} \right|_0 = a : b : c$$

Superdeformation (Axis ratio 2:1)

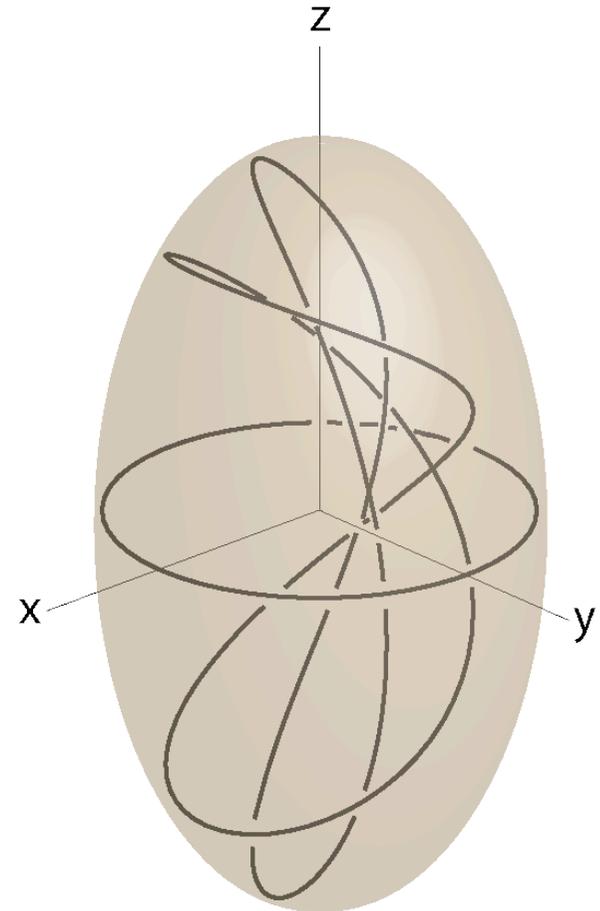


$$a : b : c = 2 : 2 : 1$$

*Bending figure of eight*



SSB to Bending (Banana) Shape



# Bending-shape phase transition in open-shell SD states

T.N., S.M., K.M., Prog. Theor. Phys. **87** (1992) 607.

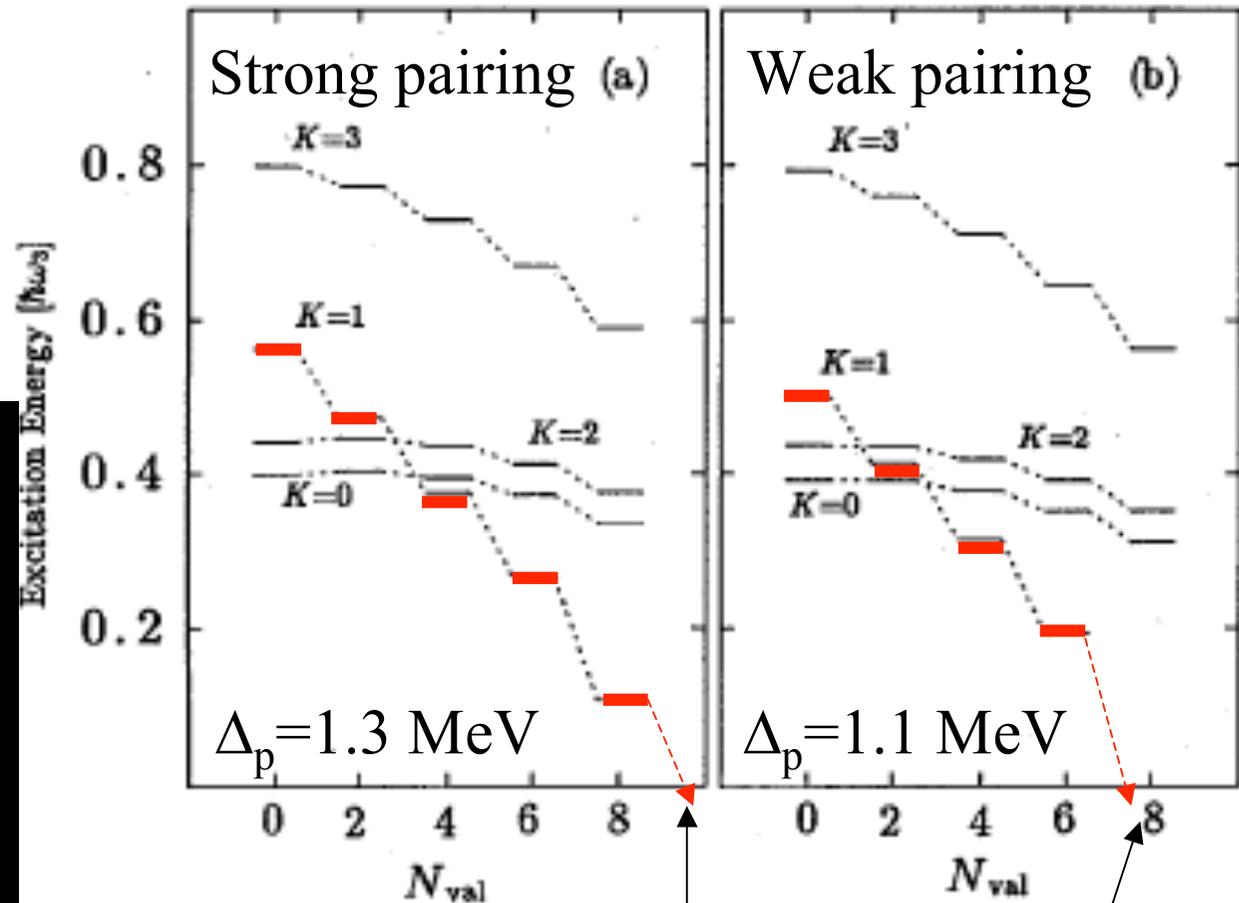
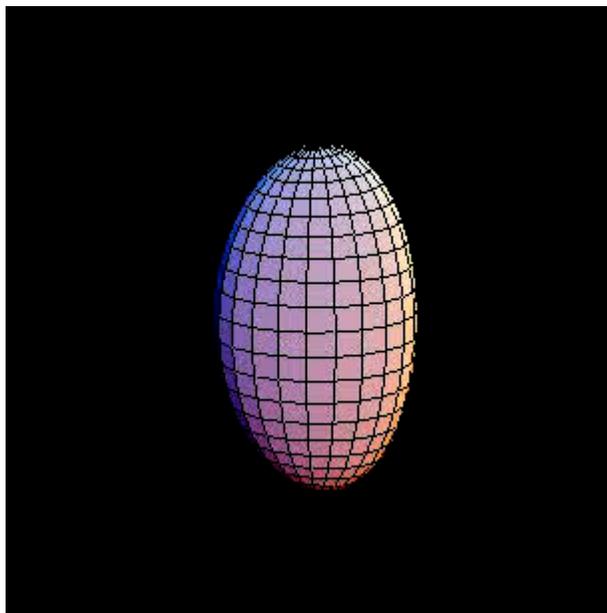
$Z=80+2$ ,  $N=80+N_{\text{val}}$

HO+QRPA

Increase of valence neutrons



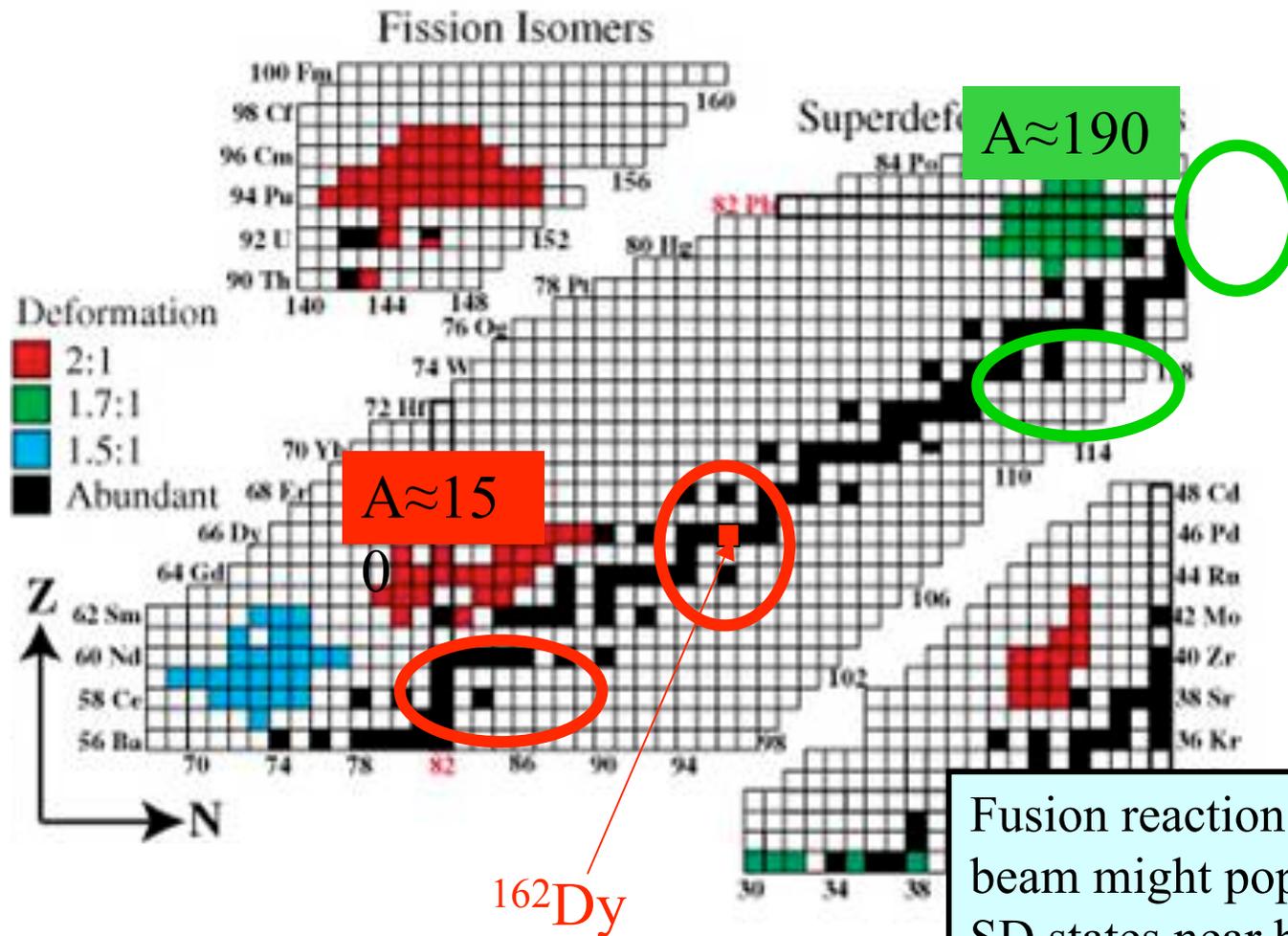
*Banana-super-deformation*



SSB towards banana shape

# Where are they?

Increasing (decreasing) valence neutrons (protons) by 8-10 leads to regions near beta-stable line



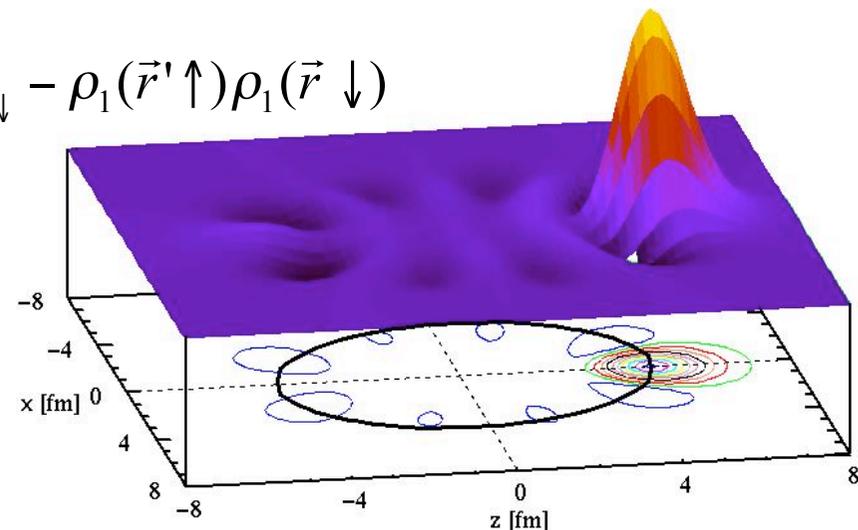
Fusion reaction with radioactive beam might populate these high-spin SD states near beta-stable line.

# From BCS to BEC ?

- Correlation becomes prominent at low density.
  - In normal nuclear density, the size of nucleonic Cooper pair is larger than the size of nucleus. ( $\sim 10$  fm or larger)
  - At low density, the pair wave function may be much more compact ( $\sim 2$  fm)

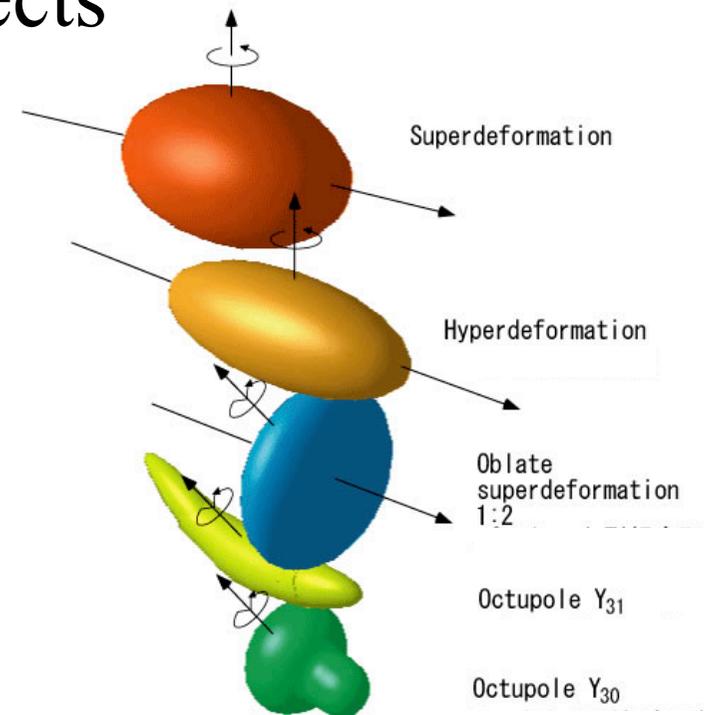
$$\rho_2^{corr}(\vec{r}' \uparrow; \vec{r} \downarrow) = \sum_{i \neq j} \delta(\vec{r} - \vec{r}_i) \delta_{\sigma_i \uparrow} \delta(\vec{r}' - \vec{r}_j) \delta_{\sigma_j \downarrow} - \rho_1(\vec{r}' \uparrow) \rho_1(\vec{r} \downarrow)$$
$$\approx |\Psi_{pair}(\vec{r} \uparrow, \vec{r}' \downarrow)|^2$$

M. Matsuo (Niigata Univ.)



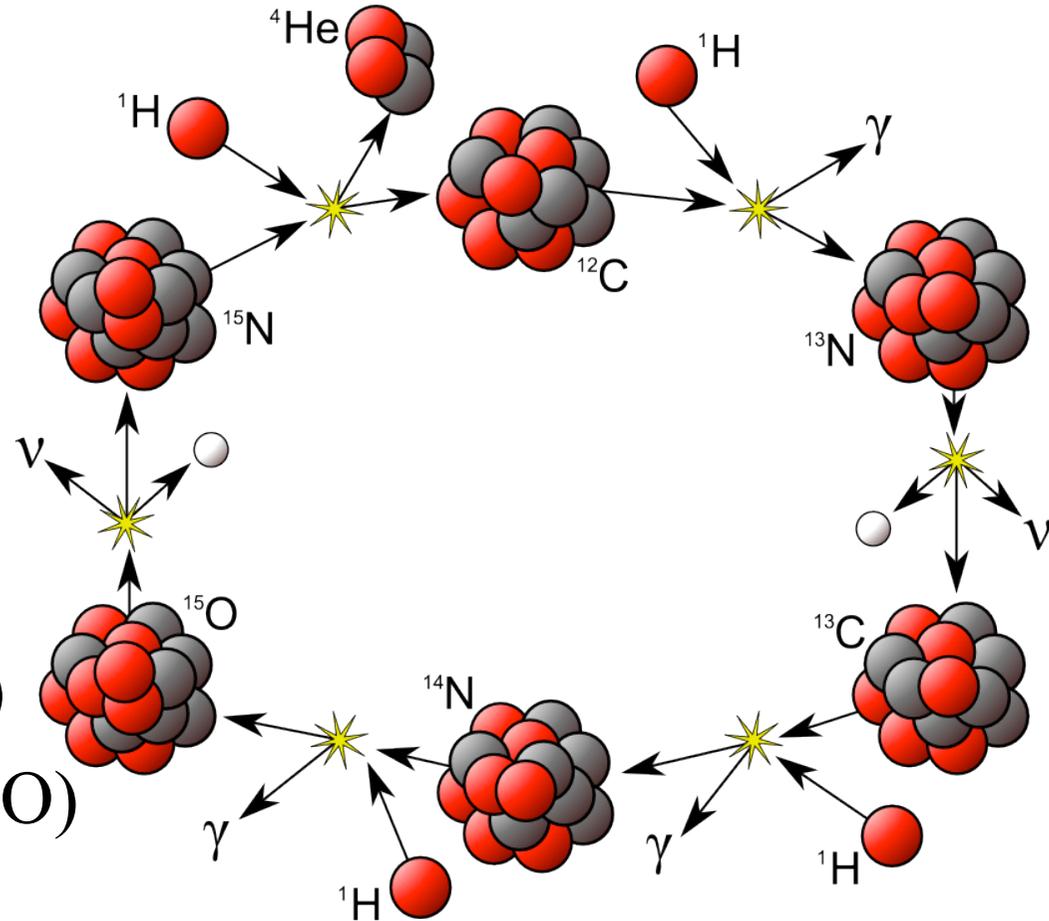
# Summary

- Nucleus is a wonderful laboratory of finite many-body systems of quantum liquid
- Nucleus shows different aspects in different time scales
- A variety of spontaneous symmetry breakings are observed in nuclei
- A new type of shape phase transition is predicted in superdeformed nuclei



# CNO Cycle

- Slowest reaction determines the energy-production efficiency
  - $p+p \rightarrow d$  (pp chain)
  - Proton capture (CNO)



	Proton	$\gamma$	Gamma Ray
	Neutron	$\nu$	Neutrino
	Positron		